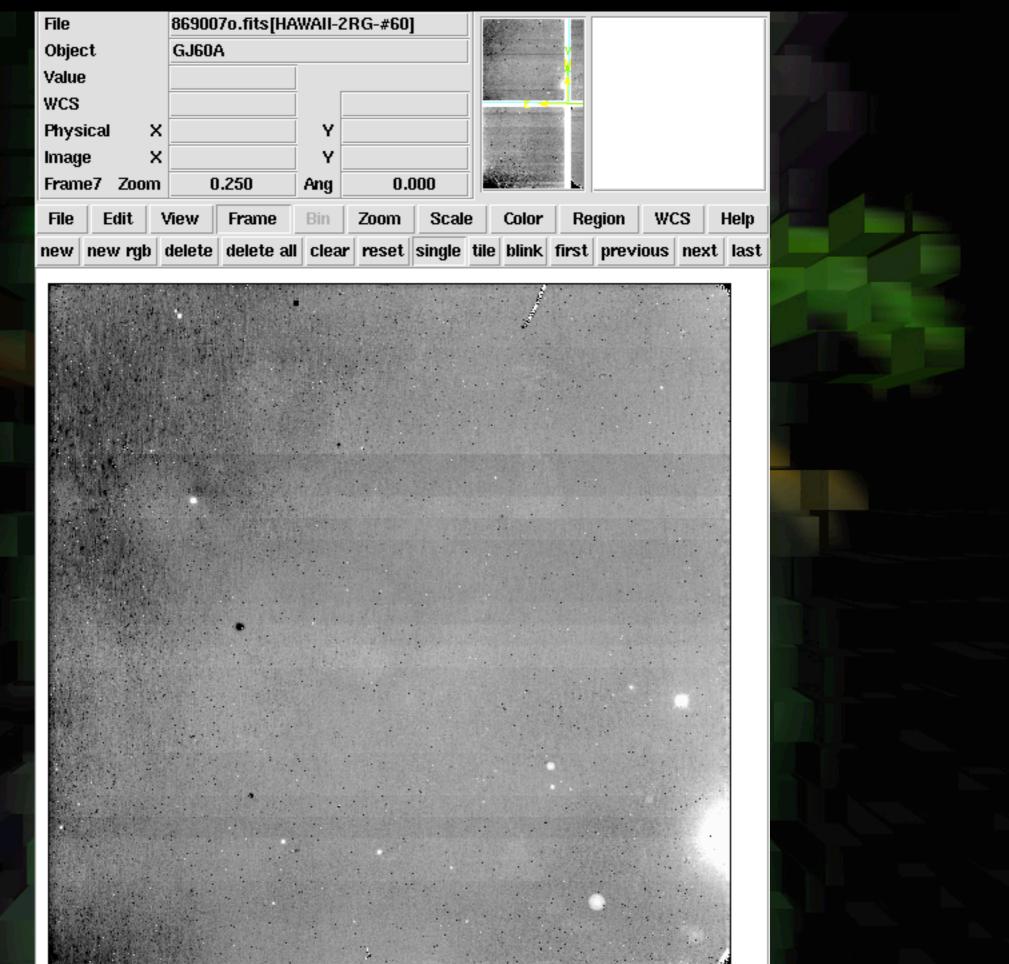


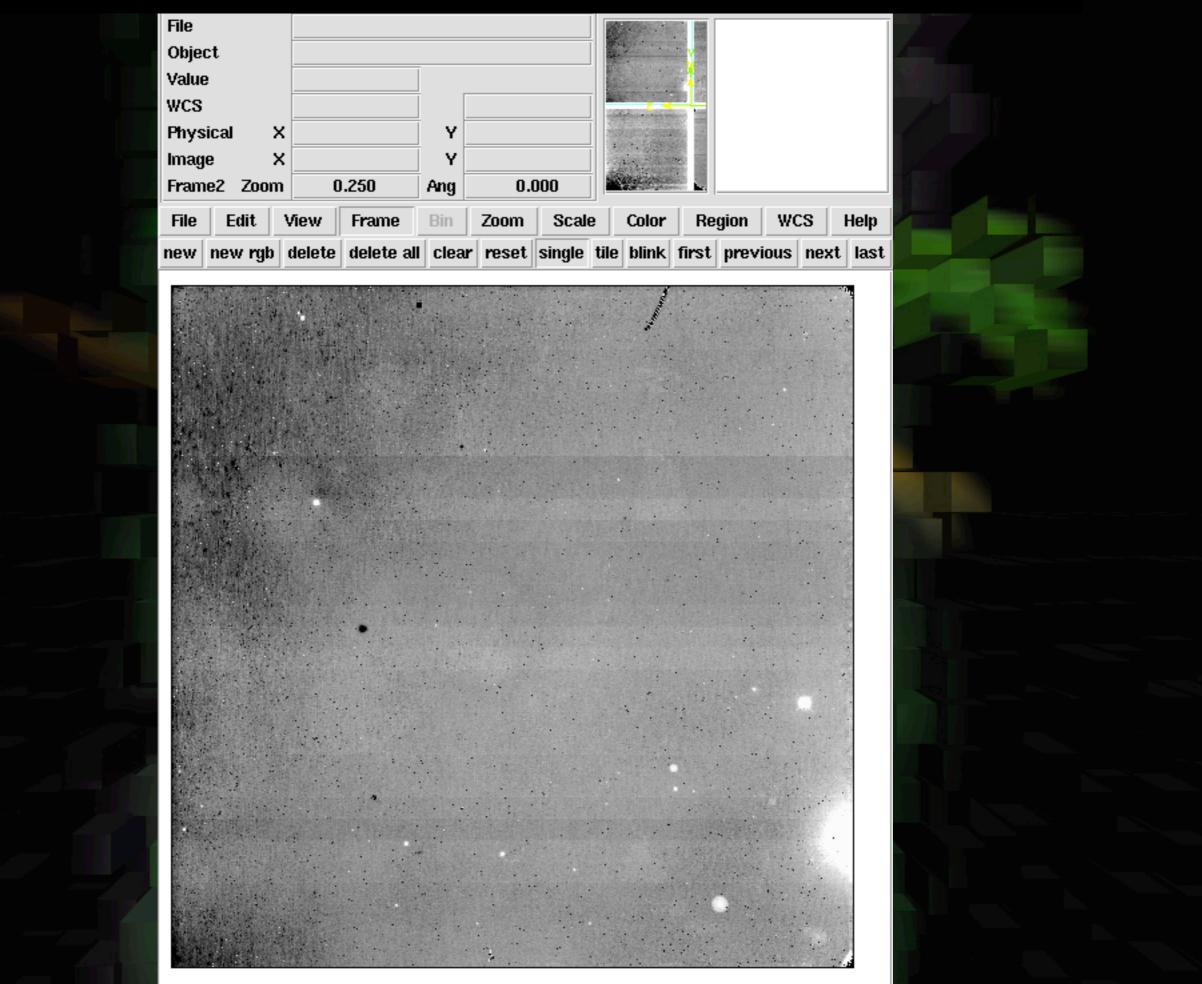
### Image Processing Recipe

- 1. Non-linearity correction (to implement)
- 2. Reference pixels correction
- 3. Dark subtraction
- 4. Bad pixel masking
- 5. Flat fielding
- 6. Guide window crosstalk masking
- 7. Sky construction/subtraction (improvement ongoing)
- 8. Common to 32 amplifiers crosstalk removal (improvement ongoing)
- 9. Positive crosstalk removal (to do)
- 10. Source extraction, IQ evaluation
- 11. Astrometry to 1" precision (linear solution per chip)
- 12. Absorption measurement based on the 2MASS catalogue

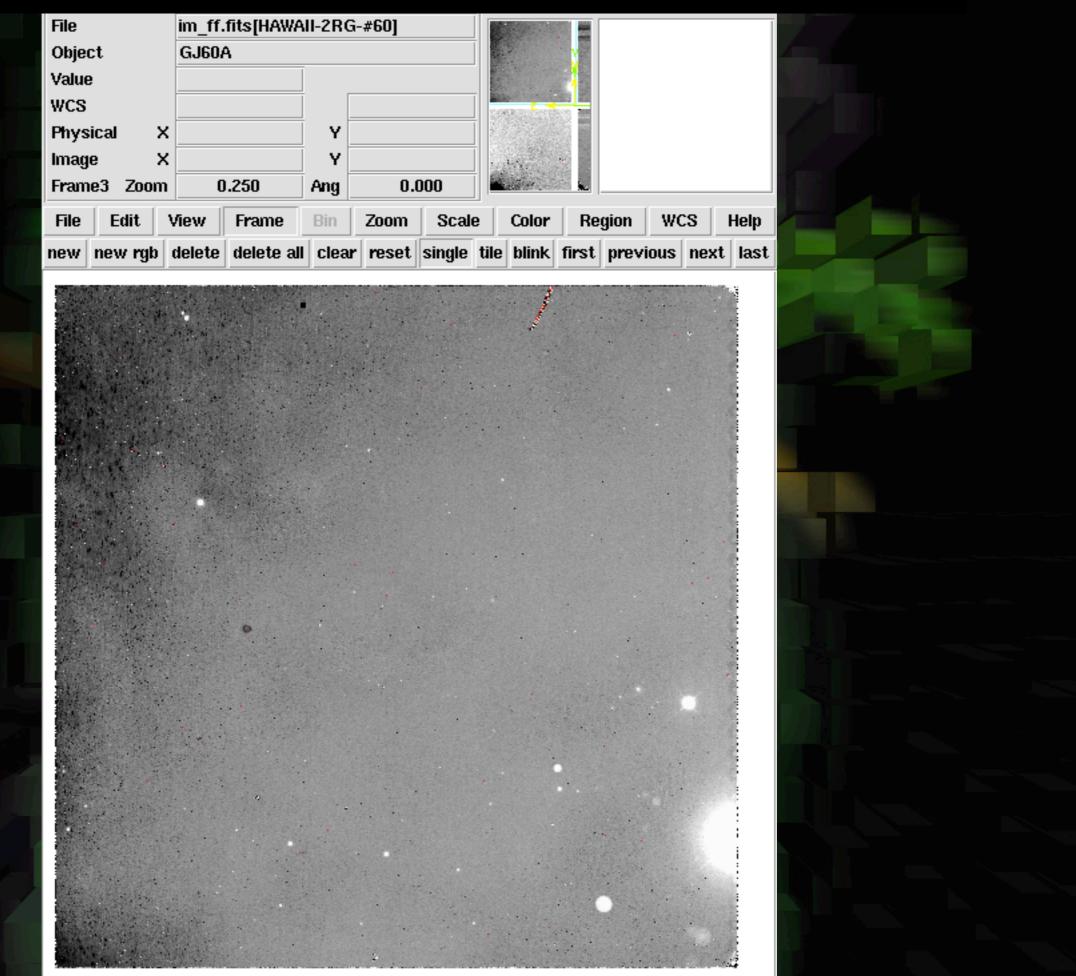
Raw



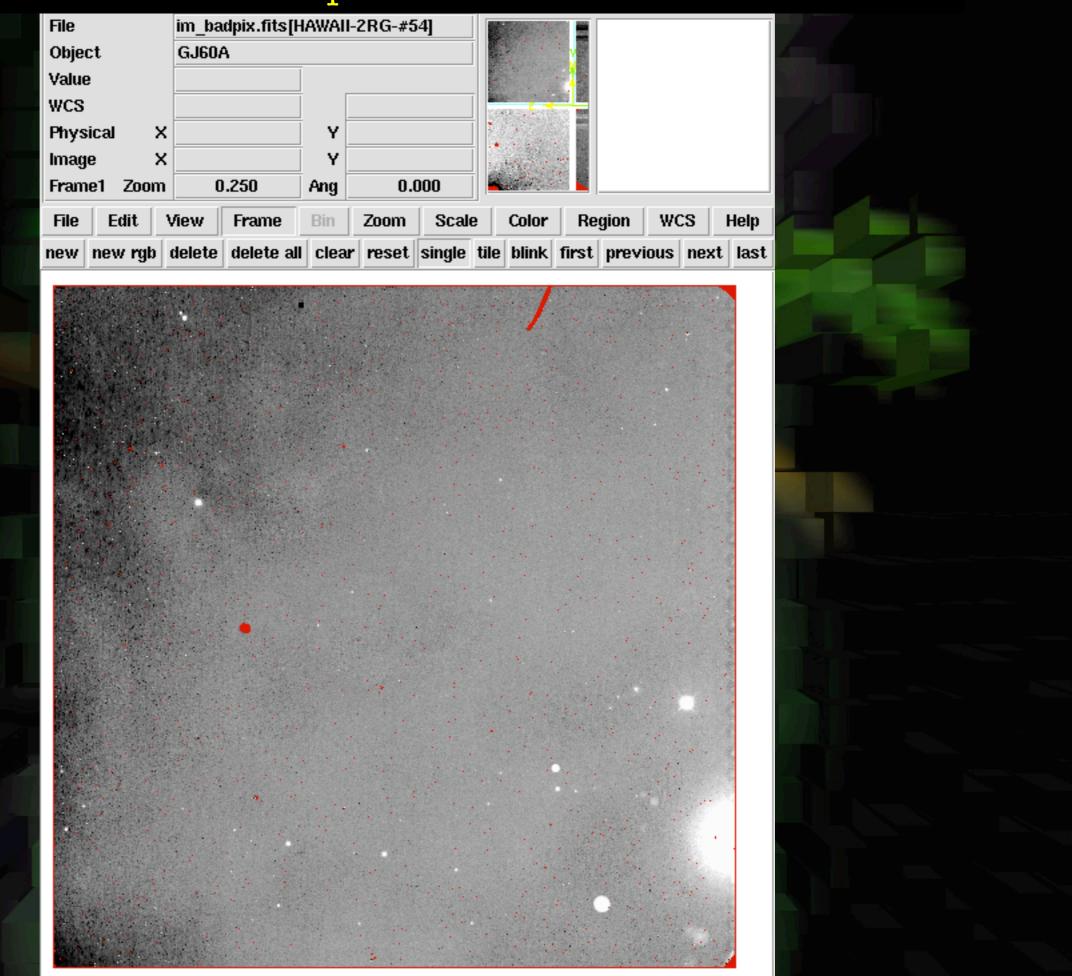
### dark subtracted



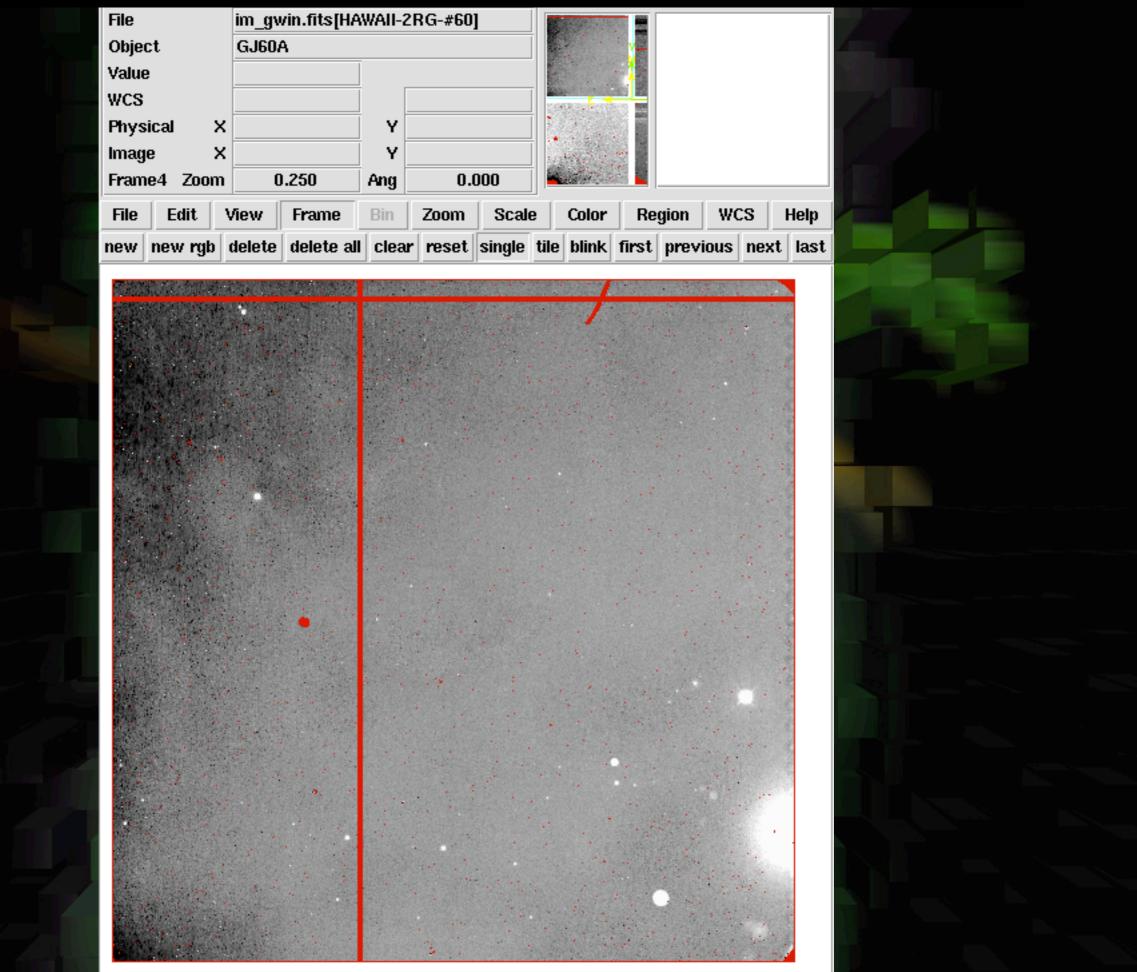
### flat fielded



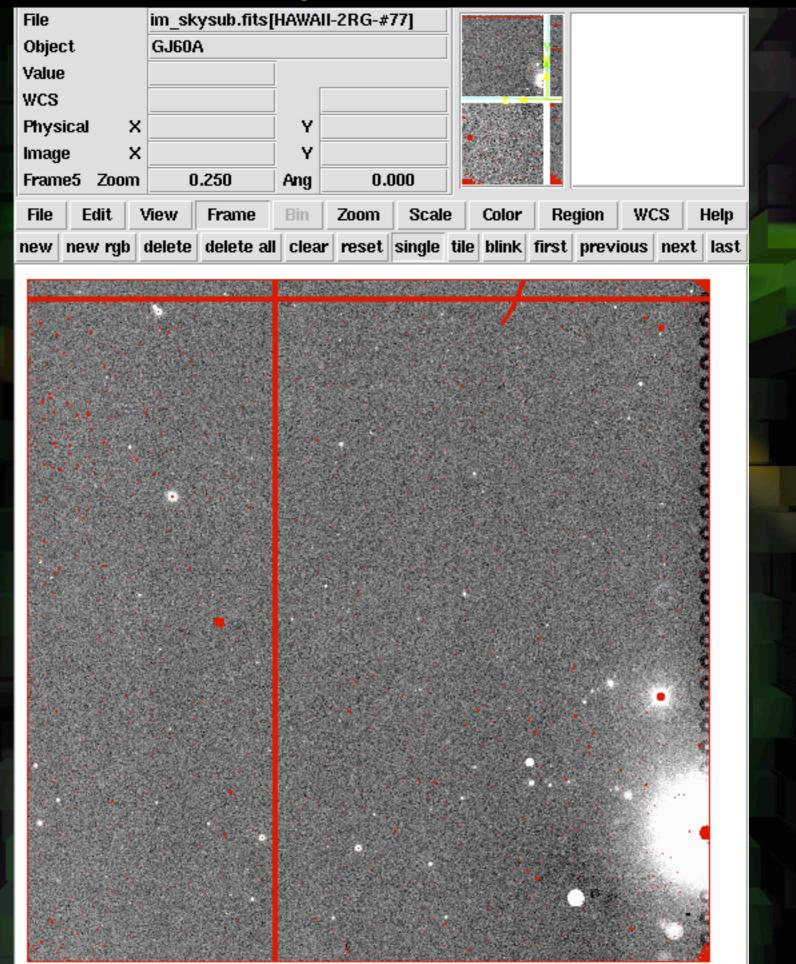
### bad pixels masked



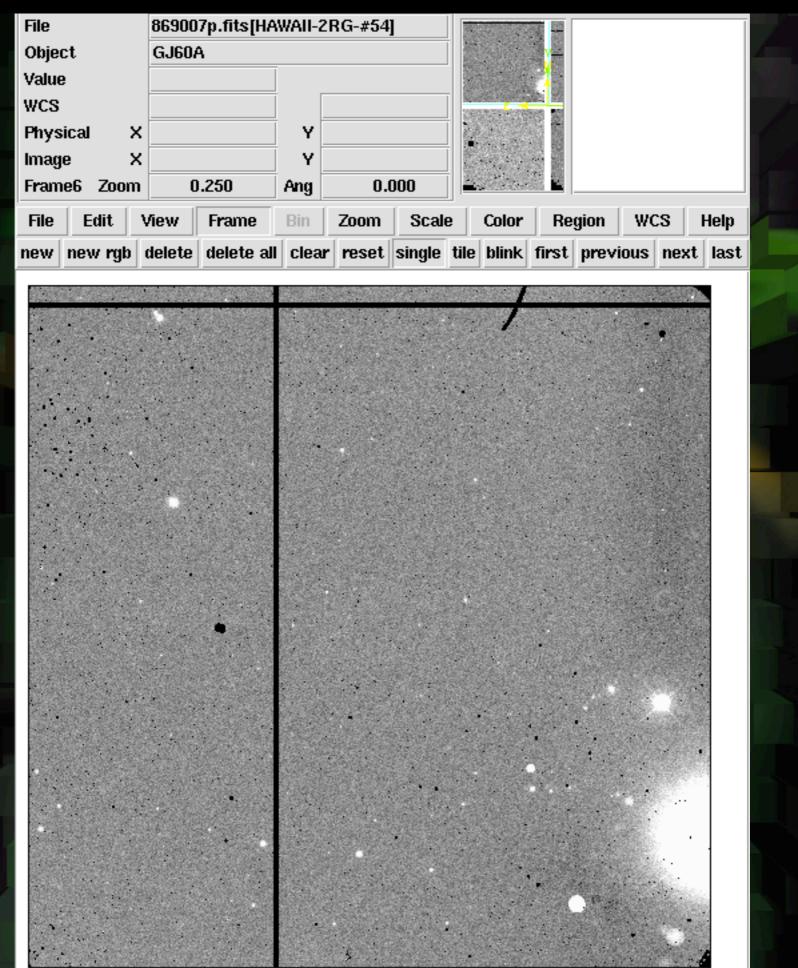
### guide window masked



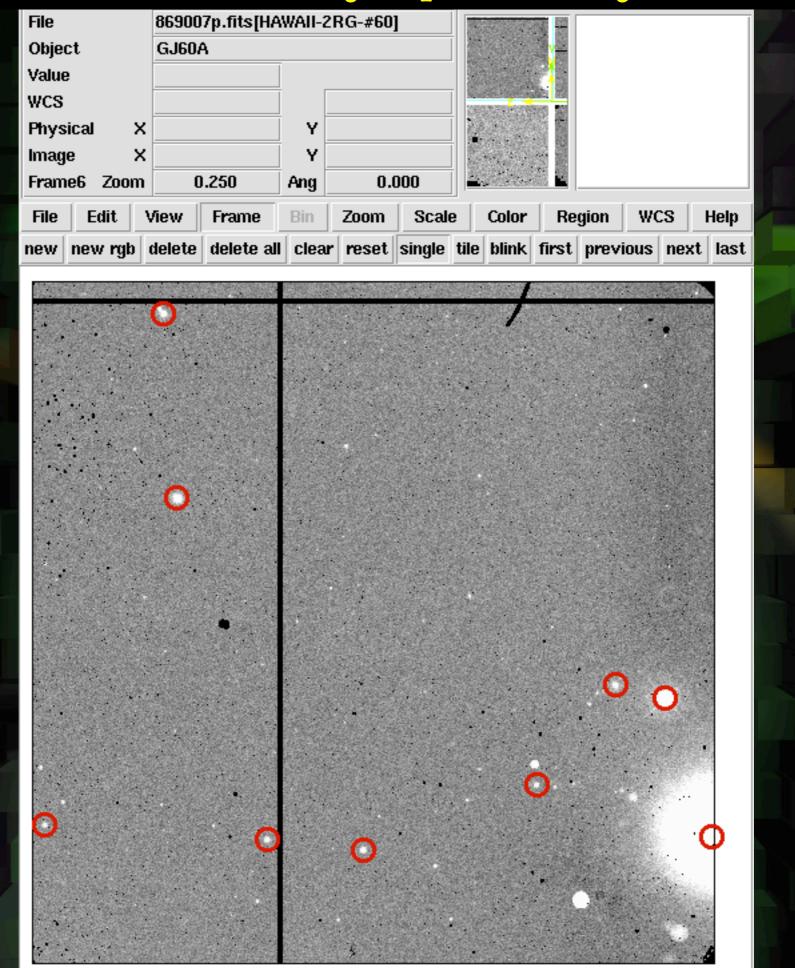
sky subtracted



### crosstalk corrected



### astrometry & photometry



Current Work Data Flow & Sky Construction Non-Linearity Crosstalk Photometric Calibration

"Sky intensity varies by 10% in 10 minutes..."

-canonical rule

Sky Construction - sliding median with source masking



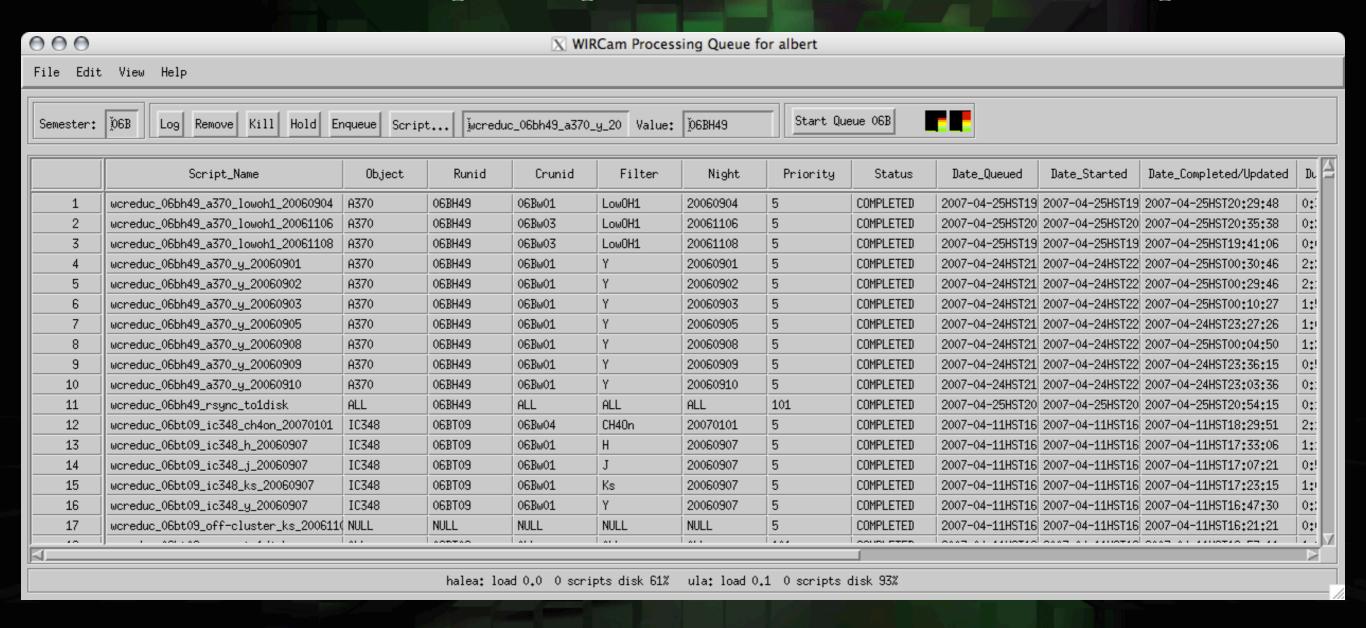
### Adjustable constraints in time and number of DPs:

example: use images taken no more than 15 minutes away and with sky positions different by at least 15"

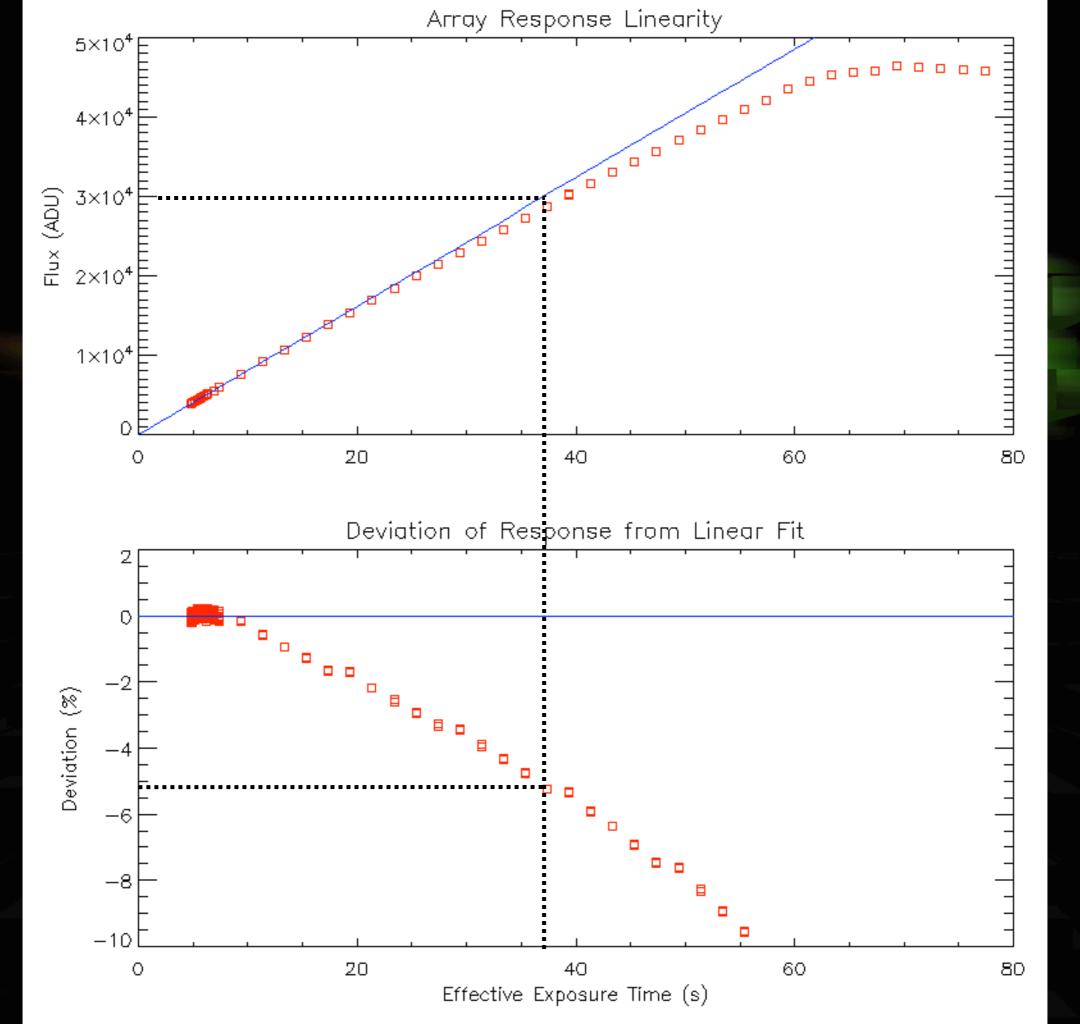
# Pipeline philosophy: Subdivide processing of images in natural groups of runid / night / target / filter

Runid	Target	Filter	Night	IDL Script		
06BH49	A370	LowOHI	20060904	wcreduc_06bh49_a370_lowoh1_20060904.pro		
		LowOHI	20061003	wcreduc_06bh49_a370_lowoh1_20061003.pro		
		Y	20060909	wcreduc_06bh49_a370_y_20060909.pro		
06BC24	D3		20070126	wcreduc_06bc24_d3_j_20070126.pro		
		Ks	20070128	wcreduc_06bc24_d3_ks_20070128.pro		
etc						

### `I`iwi runs IDL scripts in parallel on 2+ different computers

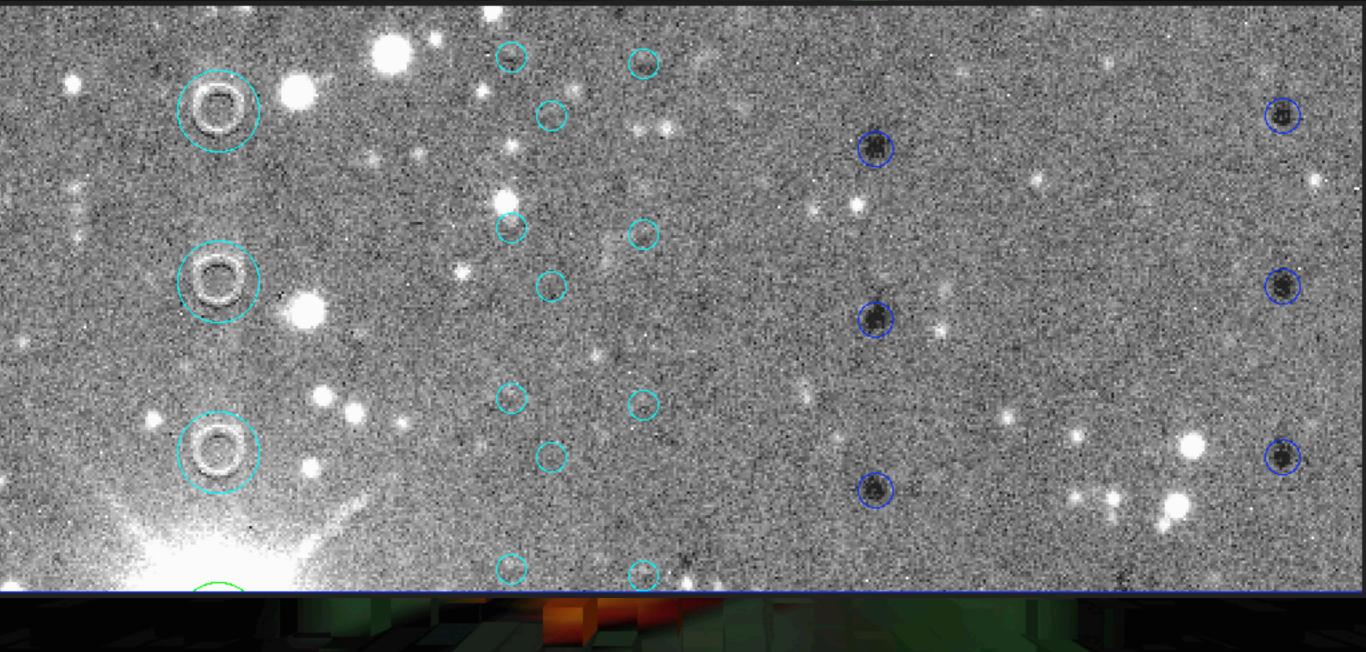


Current Work Data Flow & Sky Construction Non-Linearity Crosstalk Photometric Calibration



Current Work Data Flow & Sky Construction Non-Linearity Crosstalk Photometric Calibration

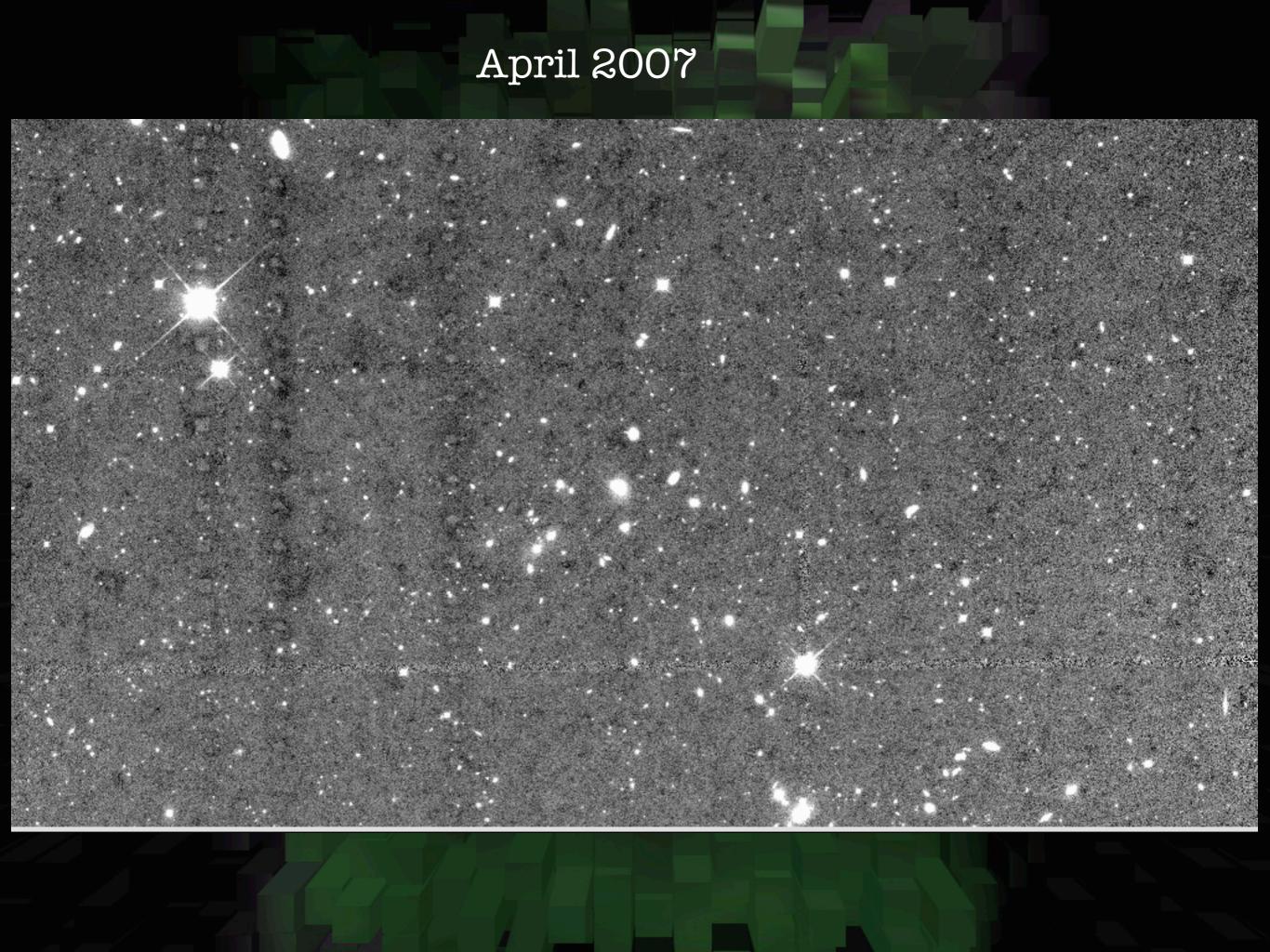
### Negative and Edge Crosstalks



Median of the 32 amplifiers isolates commonalities



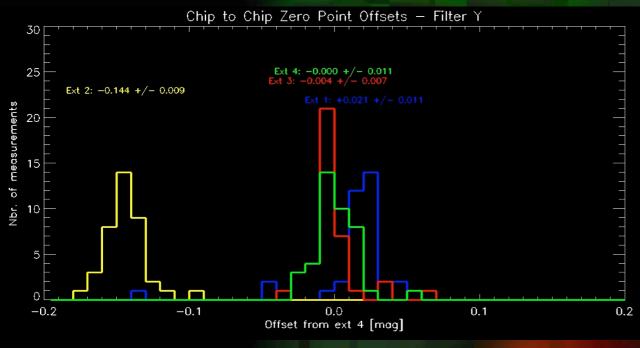
# February 2007

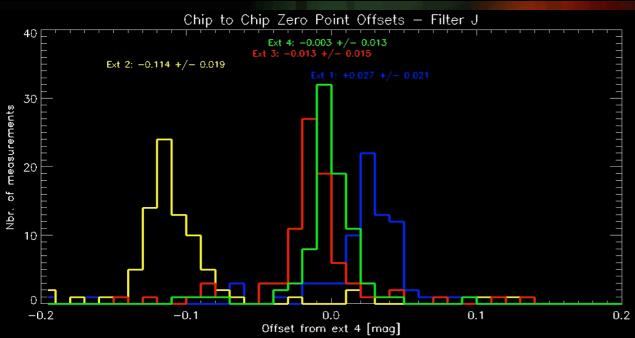


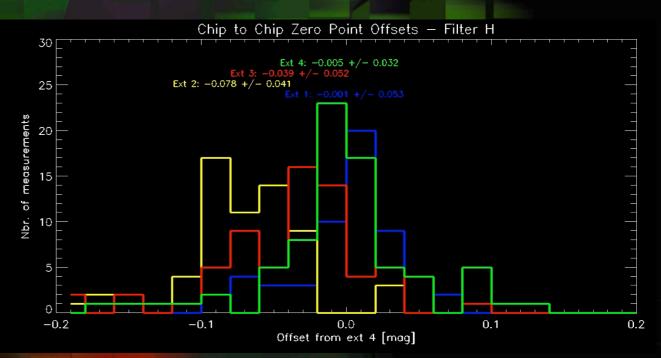
Current Work Data Flow & Sky Construction Non-Linearity Crosstalk Photometric Calibration

## Photometric Calibration on Standard Stars

### 1. Chip to Chip Zero Point Offset

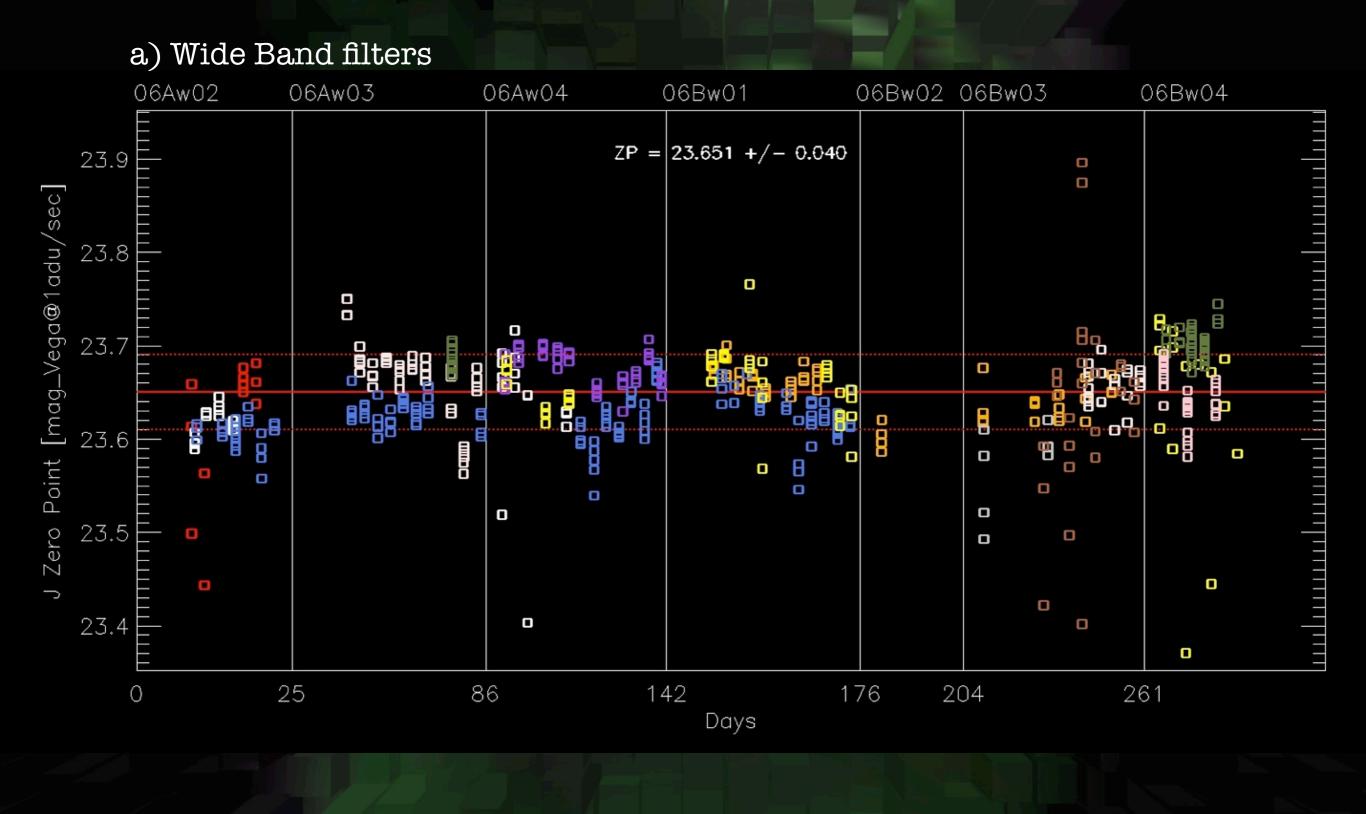






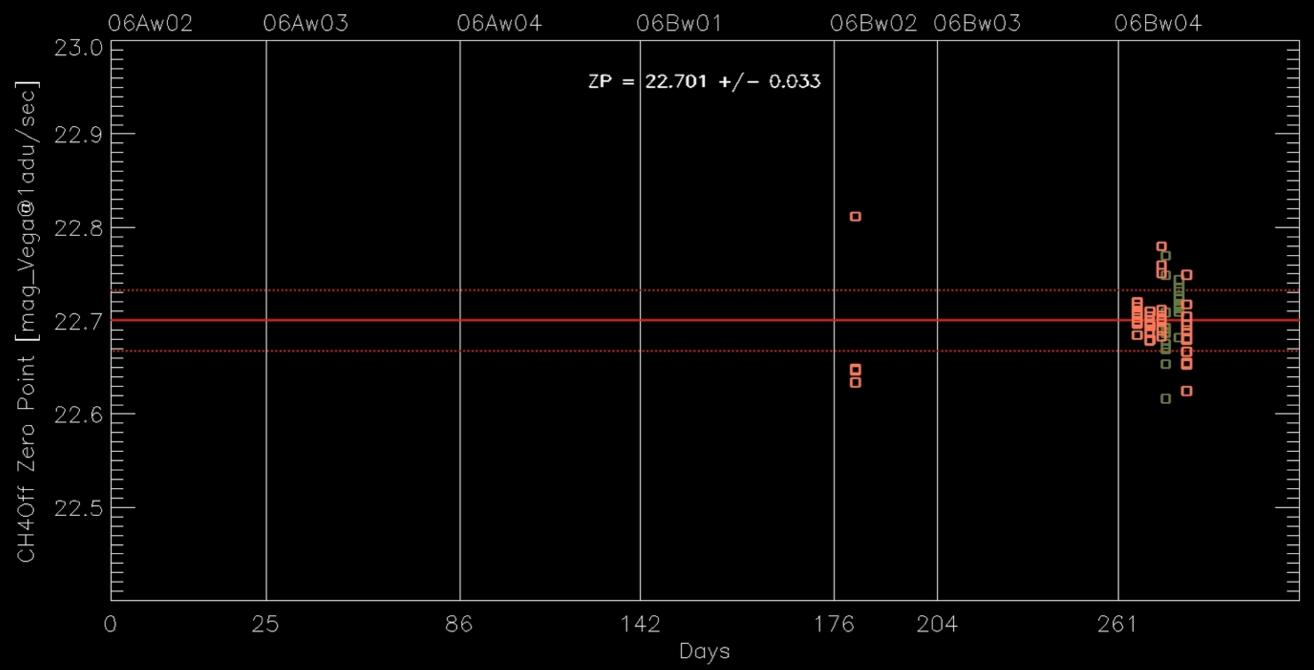


2. Zero Point vs. Time



### 2. Zero Point vs. Time





- Use 4 CALSPEC spectrophotometric standards with NIR models.
- Compute Absolute ZP using filter curves.

### 2. Zero Point vs. Time

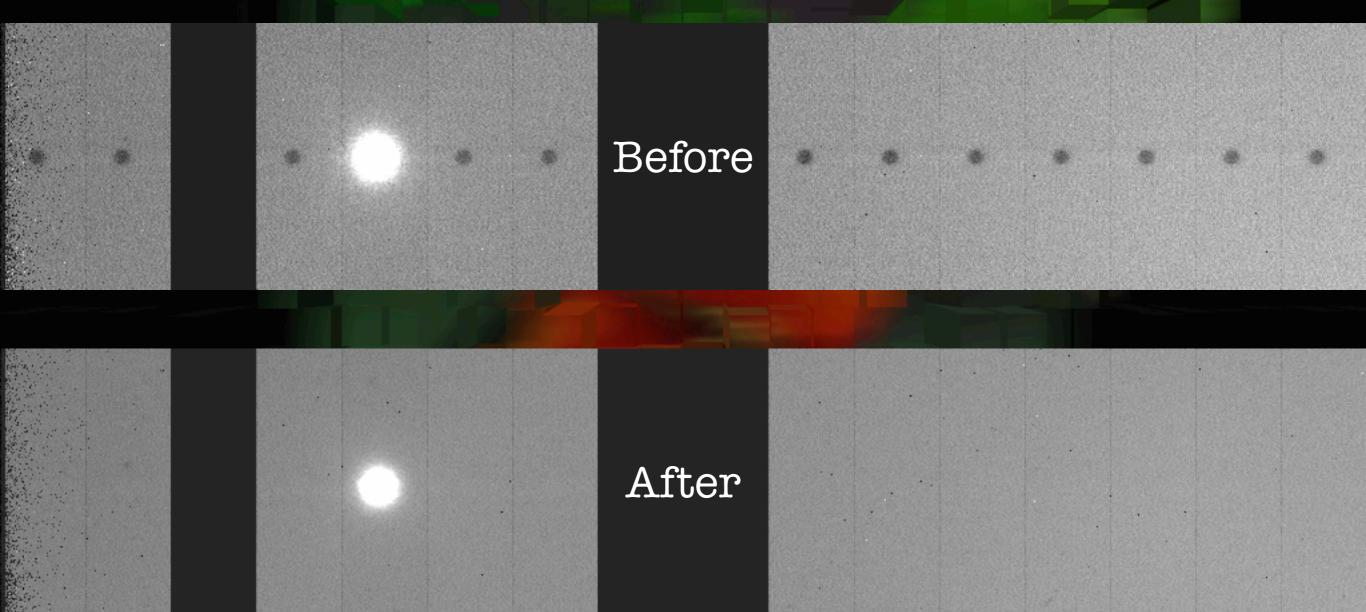
### c) Zero Point Table

Filter	ZP (mag@)	l adu/sec)
Y	23.820	0.099
J	23.651	0.040
H	23.811	0.063
KS	23.083	0.049
LOWOH1	21.008	0.072
LOWOH2	25.0	9.999
CH4ON	22.698	0.053
CH40FF	22.687	0.044
H2	20.530	0.078
KCONT	25.0	9.999
BRG	20.379	0.079

To do: Same analysis with 2MASS field stars from archived data

### Latest news! Tests in the lab

- => Engineering detector
- => 32 amps / 4 video board controller same clocking
- => Change V<sub>ref</sub> bias voltage



### Questions?

Also see 2 posters on the WIRCam instrument:

- 1 The WIRCam On-Chip Guider
- 2 WIRCam Design and Performances



the 'I'iwi Interpretor of the Wircam Images